Fusion using time-dependent density-constrained DFT

R. Keser^{1,2,a}, A.S. Umar¹, V.E. Oberacker¹, J.A. Maruhn³, and P.-G. Reinhard⁴

¹ Physics and Astronomy, Vanderbilt University, Nashville, TN 37235, USA

²RTE University, Science and Arts Faculty, Department of Physics, 53100, Rize, TURKEY

³ Institut für Theoretische Physik, Goethe-Universität, D-60438 Frankfurt am Main, Germany

⁴ Institut für Theoretische Physik, Universität Erlangen, D-91054 Erlangen, Germany

Abstract. We present results for calculating fusion cross-sections using a new microscopic approach based on a time-dependent density-constrained DFT calculations. The theory is implemented by using densities and other information obtained from TDDFT time-evolution of the nuclear system as a constraint on the density for DFT calculations.

1 Introduction

The investigation of internuclear potentials for heavy-ion collisions is of fundamental importance for the study of fusion reactions as well as for the formation of superheavy elements and nuclei far from stability. Recently, we have developed a new method to extract ion-ion interaction potentials directly from the time-dependent Hartree-Fock (TDHF) time-evolution of the nuclear system [1]. In the density-constrained TDHF (DC-TDHF) approach the TDHF time-evolution takes place with no restrictions. At certain times during the evolution the instantaneous density is used to perform a static Hartree-Fock minimization while holding the neutron and proton densities constrained to be the corresponding instantaneous TDHF densities. In essence, this provides us with the TDHF dynamical path in relation to the multi-dimensional static energy surface of the combined nuclear system. In this approach there is no need to introduce constraining operators which assume that the collective motion is confined to the constrained phase space. In short, we have a self-organizing system which selects its evolutionary path by itself following the microscopic dynamics. Some of the effects naturally included in the DC-TDHF calculations are: neck formation, mass exchange, internal excitations, deformation effects to all order, as well as the effect of nuclear alignment for deformed systems. The DC-TDHF theory provides a comprehensive approach to calculating fusion barriers in the mean-field limit. The fusion cross-sections are subsequently calculated by integrating the Schrödinger equation using the incoming-wave boundary condition method (IWBC) as outlined in Ref. [2]. The theory has been applied to calculate fusion cross-sections for a large number of systems [2–8]. In this contribution we will outline the DC-TDHF method and give new examples of its application to the calculation of fusion cross-sections for several systems.

^ae-mail: recep.keser@erdogan.edu.tr

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2 Recent Applications of the DC-TDHF Method

In this Section we give some recent examples of DC-TDHF calculations of heavy-ion potentials and cross-sections. Recently, we have studied the fusion of very neutron rich light nuclei that may be important to determine the composition and heating of the crust of accreting neutron stars [11]. The main focus was the O+O and C+O systems. For the ${}^{16}O+{}^{16}O$ system we have shown excellent agreement between our calculations and the low energy data from Refs. [12, 13]. We have also extended this work to higher energies to see how our results compare with the available data. The reactions of light systems at high energies (2 – 3 times the barrier height) is complicated both experimentally and theoretically due to the presence of many breakup channels and excitations. All the data we could find date back to late 1970's [14–16]. Experimental findings differ considerably in this energy regime as can be seen from Fig. 1. Recent analysis of the ${}^{16}O+{}^{16}O$ system system by H. Esbensen [17] primarily uses the data of Tserruya *et al.* [15]. We expect the TDHF results to yield a higher fusion cross-section since many of the breakup channels are not naturally available in TDHF. However, a close investigation of the TDHF dynamics and the microscopically calculated excitation energy clearly indicate that a significant portion of the collective kinetic energy is not equilibriated. It may be plausible to con-

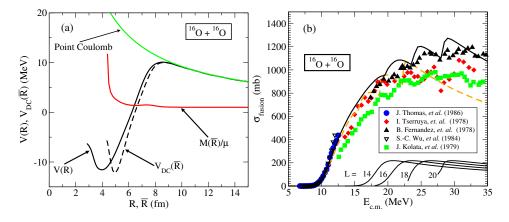


Figure 1. (a) Ion-Ion potential and effective mass for ¹⁶O+¹⁶O. (b) Corresponding fusion cross-sections.

sider the direct influence of the excitation energy, $E^*(R)$, on the fusion barriers by making an analogy with the coupled-channel approach and construct a new potential $V^*(R) = V(R) + E^*(R)$, which has all the excitations added to the ion-ion potential V(R) that should be calculated at higher energies to minimize the nuclear rearrangements (frozen-density limit). The resulting potentials somewhat resemble the repulsive-core coupled-channel potentials of Ref. [18]. This approach does lead to improvements in cases where most of the excitation energy is in the form of collective excitations rather than irreversible stochastic dissipation (true especially for lighter systems). The viability of this approach requires further examination and will be studied in the future. It is interesting to note that the gross oscillations in the cross-section at higher energies are correctly reproduced in our calculations. This is simply due to opening of new *L*-channels as we increase the collision energy. Individual contributions to the cross-section from higher *L* vales are also shown on the lower part of the plot.

In Fig. 2a we show the DC-TDHF potential barriers for the C+O system. The higher barrier corresponds to the ${}^{12}C+{}^{16}O$ system and has a peak energy of 7.77 MeV. The barrier for the ${}^{12}C+{}^{24}O$ system occurs at a slightly larger *R* value with a barrier peak of 6.64 MeV. Figure 2b shows the corresponding cross sections for the two reactions. Also shown are the experimental data from

INPC 2013

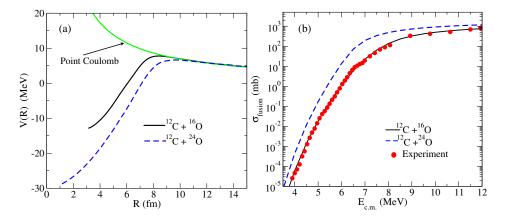


Figure 2. (a) Ion-Ion potential for various isotopes of the C+O system. (b) Corresponding cross-sections.

Refs. [19–21]. The DC-TDHF potential reproduces the experimental cross-sections quite well for the ${}^{12}C+{}^{16}O$ system, and the cross section for the neutron rich ${}^{12}C+{}^{24}O$ is predicted to be larger than that for ${}^{12}C+{}^{16}O$.

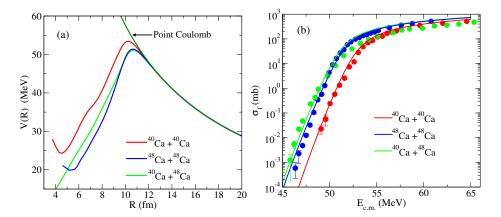


Figure 3. (a) Ion-Ion potential for various isotopes of the Ca+Ca system. (b) Corresponding cross-sections.

Figures 3a and 3b show the corresponding potentials and cross-sections for the Ca+Ca system [22], which was the subject of recent experimental studies [23]. The observed trend for subbarrier energies is typical for DC-TDHF calculations when the underlying microscopic interaction gives a good representation of the participating nuclei. Namely, the potential barrier corresponding to the lowest collision energy gives the best fit to the sub-barrier cross-sections since this is the one that allows for more rearrangements to take place and grows the inner part of the barrier. Considering the fact that historically the low-energy sub-barrier cross-sections of the ⁴⁰Ca+⁴⁸Ca system have been the ones not reproduced well by the standard models, the DC-TDHF results are quite satisfactory, indicating that the dynamical evolution of the nuclear density in TDHF gives a good overall description of the collision process. The shift of the cross-section curve with increasing collision energy is typical.

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In principle one could perform a DC-TDHF calculation at each energy above the barrier and use that cross-section for that energy. However, this would make the computations extremely time consuming and may not provide much more insight. The trend at higher energies for the ⁴⁰Ca+⁴⁸Ca system is atypical. The calculated cross-sections are larger than the experimental ones by about a factor of two. Such lowering of fusion cross-sections with increasing collision energy is commonly seen in lighter systems where various inelastic channels, clustering, and molecular formations are believed to be the contributing factors.

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References

- [1] A. S. Umar and V. E. Oberacker Phys. Rev. C 74 021601(R) (2006).
- [2] A. S. Umar and V. E. Oberacker Eur. Phys. J. A 39, 243 (2009).
- [3] A. S. Umar and V. E. Oberacker Phys. Rev. C 74 024606 (2006).
- [4] A. S. Umar and V. E. Oberacker Phys. Rev. C 74 061601(R) (2006).
- [5] A. S. Umar and V. E. Oberacker Phys. Rev. C 76 014614 (2007).
- [6] A. S. Umar and V. E. Oberacker Phys. Rev. C 77 064605 (2008).
- [7] A. S. Umar, V. E. Oberacker, J. A. Maruhn and P.-G. Reinhard, Phys. Rev. C 81, 064607 (2010).
- [8] V. E. Oberacker, A. S. Umar, J. A. Maruhn, and P.-G. Reinhard, Phys. Rev. C 82, 034603 (2010).
- [9] A. S. Umar, M. R. Strayer, R. Y. Cusson, P. -G. Reinhard, and D. A. Bromley, Phys. Rev. C 32, 172 (1985).
- [10] R. Y. Cusson, P.-G. Reinhard, M. R. Strayer, J. A. Maruhn, and W. Greiner, Z. Phys. A 320, 475 (1985).
- [11] A. S. Umar, V. E. Oberacker, and C. J. Horowitz, Phys. Rev. C 85, 055801 (2012).
- [12] J. Thomas, Y. T. Chen, S. Hinds, K. Langanke, D. Meredith, M. Olson, and C. A. Barnes, Phys. Rev. C 31, 1980 (1985).
- [13] J. Thomas, Y. T. Chen, S. Hinds, D. Meredith, and M. Olson, Phys. Rev. C 33, 1679 (1986).
- [14] B. Fernandez, C. Gaarde, J. S. Larsen, S. Pontoppidan, and F. Videbaek, Nucl. Phys. A306, 259 (1978).
- [15] I. Tserruya, Y. Risen, D. Pelte, A. Gavron, H. Oeschler, D. Berndt, and H.L Harney, *Phys. Rev.* C 18, 1688 (1978).
- [16] J. J. Kolata, R. M. Freeman, F. Haas, B. Heusch, and A. Gallmann, Phys. Rev. C 19, 2237 (1979)
- [17] H. Esbensen, Phys. Rev. C 77, 054608 (2008).
- [18] H. Esbensen et al., Phys. Rev. C 82, 054621 (2010).
- [19] C. L. Jiang, K. E. Rehm, B. B. Back, and R. V. F. Janssens, Phys. Rev. C 75, 015803 (2007).
- [20] H. W. Becker et al., Z. Phys. A303, 305 (1981).
- [21] L. R. Gasques et al., Phys. Rev. C 72, 025806 (2005).
- [22] R. Keser, A. S. Umar, and V. E. Oberacker, Phys. Rev. C 85, 044606 (2012).
- [23] G. Montagnoli et al., Phys. Rev. C 85, 024607 (2012).