Enhancements of high order cumulants across the 1st order phase transition boundary

Lijia Jiang^{1,2,*}, Shanjin Wu^{2,**}, and Huichao Song^{2,3,4,***}

Abstract. In this proceeding, we investigate the dynamical evolution of the σ field with a trajectory across the 1st order phase transition boundary, using the Langevin equation from the linear sigma model. We find the high order cumulants of the σ field are largely enhanced during the dynamical evolution, compared with the equilibrium values, due to the supercooling effect of the first order phase transition.

1 Introduction

The STAR Beam Energy Scan (BES) program aims at searching and locating the QCD critical point. It was proposed that the fluctuations of net baryon number and net protons are largely enhanced in the critical region due to the increased correlation length [1, 2]. In experiment, large deviations of high order cumulants and cumulants ratios of net protons from the Poisson baselines have been observed at the collision energies below 40 GeV [3–5], indicating the potential of discovering the critical point.

Many past research has studied the static critical fluctuations near the critical point, which could explain the acceptance dependence of the net proton's fluctuations, but could not qualitatively describe all the cumulants measured in experiment [6–8]. Recently, the investigations on the dynamical critical fluctuations based on the Fokker-Plank equation and Langevin dynamics have shown that critical slowing-down effects play important roles in the critical regime, where the signs of high order dynamical cumulants of the σ field are even different from the corresponding equilibrium values [9, 10].

In this proceeding, we focus on studying the dynamical evolution of the σ 's cumulants with the 1st order phase transition scenario, using the Langevin dynamics within the framework of the linear sigma model. We will show that the high order cumulants of the σ field are largely enhanced during the dynamical evolution when compared with the equilibrium values, due to the supercooling effect across the first order phase transition boundary.

¹ Frankfurt Institute for Advanced Studies, Ruth Moufang Strasse 1, D-60438, Frankfurt am Main, Germany

²Department of Physics and State Key Laboratory of Nuclear Physics and Technology, Peking University, Beijing 100871, China

³Collaborative Innovation Center of Quantum Matter, Beijing 100871, China

⁴Center for High Energy Physics, Peking University, Beijing 100871, China

^{*}jiang@fias.uni-frankfurt.de

^{**}shanjinwu2014@pku.edu.cn

^{***}huichaosong@pku.edu.cn

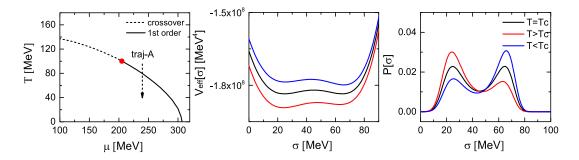


Figure 1. Left panel: the phase diagram of the linear sigma model. The middle and right panel: the effective potential $V_{eff}(\sigma)$ and probability distribution function $P[\sigma]$ at different temperature $T > T_c$, $T = T_c$ and $T < T_c$, along a chosen trajectory with $\mu = 240$ MeV.

2 The formulism

Within the framework of the linear sigma model with constituent quarks, we solve the time evolution of the σ field based on the Langevin equation. The effective potential of this model gives a phase diagram with cross-over, the first order phase transition as well as a critical point [11–13]. The long wavelength mode of the σ field is the order parameter of the chiral phase transition, and its mass vanish at the critical point. According to the classification of [14], the dynamics of the σ field in the critical regime belongs to model A, which only evolves the non-conserved order parameter field. The Langevin equation of the σ field is written as:

$$\partial^{\mu}\partial_{\mu}\sigma\left(t,x\right) + \eta\partial_{t}\sigma\left(t,x\right) + \frac{\delta V_{eff}\left(\sigma\right)}{\delta\sigma} = \xi\left(t,x\right),\tag{1}$$

which is a semi-classical equation. The damping η and noise $\xi(t, x)$ term in the above equation satisfy the fluctuation-dissipation theorem: $\langle \xi(t)\xi(t')\rangle \sim \eta \delta(t-t')$. The equilibrium thermodynamical potential is obtained by integrating out the thermal quarks:

$$V_{eff}(\sigma) = U(\sigma) + \Omega_{\bar{q}q}(\sigma) = \frac{\lambda^2}{4} \left(\sigma^2 - v^2\right)^2 - h_q \sigma - U_0 + \Omega_{\bar{q}q}(\sigma). \tag{2}$$

Here $U(\sigma)$ represents the vacuum potential of the σ field, $\Omega_{\bar{q}q}(\sigma) = -d_q \int \frac{d^3p}{(2\pi)^3} \{\omega + T \ln[1 + e^{-(\omega-\mu)/T}] + T \ln[1 + e^{-(\omega+\mu)/T}] \}$ comes from the contribution of thermal quarks, the particle energy $\omega = \sqrt{p^2 + M^2}$, and the effective mass of quarks $M(\sigma) = m_0 + g\sigma$, with the current quark mass $m_0 = 5.5$ MeV, and coupling g = 3.3. The parameters h_q , v, λ , U_0 in the effective potential are determined by the properties of vacuum hadrons. Due to the spontaneously broken of chiral symmetry in vacuum, the vacuum expectation of σ field is $\langle \sigma \rangle = f_{\pi} = 93$ MeV. h_q is the symmetry explicitly broken term, $h_q = f_{\pi} m_{\pi}^2$ with $m_{\pi} = 138$ MeV. The parameter v is determined by $v^2 = f_{\pi}^2 - m_{\pi}^2/\lambda^2$, the in vacuum mass of σ is $m_{\sigma} \sim 600$ MeV by setting $\lambda^2 = 20$. The zero-point energy $U_0 = m_{\pi}^4/(4\lambda^2) - f_{\pi}^2 m_{\pi}^2$. Note that we have neglected the fluctuations of $\vec{\pi}$, since its mass is finite in the critical regime.

From the thermodynamic potential in eq. (2), we depict the phase diagram on the μ -T plane with a critical point located at (μ_c , T_c) = (205, 100.2) MeV, as shown in Fig. 1 (left). The middle and right panel of Fig. 1 present the effective potential and probability distribution function of the σ field

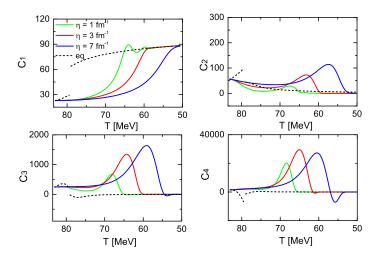


Figure 2. Dynamical evolution of the σ 's cumulants along trajectory-A across the first order phase transition with $\mu = 240$ MeV. The colored lines represent the dynamical cumulants with different damping coefficients. The black dotted lines are the corresponding equilibrium cumulants.

$$\left(P\left[\sigma\left(\mathbf{x}\right)\right] \sim \exp\left(-E\left(\sigma\right)/T\right) \text{ with } E\left(\sigma\right) = \int d^3x \left[\frac{1}{2}\left(\nabla\sigma\left(x\right)\right)^2 + V_{eff}\left(\sigma\left(x\right)\right)\right]\right)$$
 at different temperature $T > T_c$, $T = T_c$ and $T < T_c$, along a chosen trajectory with $\mu = 240$ MeV.

For the numerical implementation of the Langevin dynamics, we need to input the initial configurations of the σ field to start the simulations. Here, we construct the event-by-event initial conditions according to the probability function $P\left[\sigma\left(\mathbf{x}\right)\right]$. In addition, the simulations of the Langevin equation requires to input the changing temperature and chemical potential profiles $T\left(t,x,y,z\right)$ and $\mu\left(t,x,y,z\right)$ from the heat bath. For simplicity, we assume the system is isothermal and evolves along a trajectory across the first order phase transition boundary with a fixed chemical potential, and the temperature decreases in a Hubble like way: $\frac{T\left(t\right)}{T_{0}} = \left(\frac{t}{t_{0}}\right)^{-0.45}$. From the dynamical evolution of the σ field, one can obtain the spatial information of $\sigma(x)$ at each

From the dynamical evolution of the σ field, one can obtain the spatial information of $\sigma(x)$ at each time step for each event, and then sum all the event configurations to obtain the associated moments: $\mu'_n = \langle \sigma^n \rangle = \frac{\int d\sigma \sigma^n P[\sigma]}{\int d\sigma P[\sigma]}$, where $\sigma = \int d^3x \sigma(\mathbf{x})$. The cumulants of the σ field can be calculated from these various moments with:

$$C_1 = \mu_1', \ C_2 = \mu_2' - \mu_1'^2, \ C_3 = \mu_3' - 3\mu_2'\mu_1' + 2\mu_1'^3, \ C_4 = \mu_4' - 4\mu_3'\mu_1' - 3\mu_2'^2 + 12\mu_2'\mu_1'^2 - 6\mu_1'^4. \ \ (3)$$

3 Numerical results

In this proceeding, we focus on investigating the evolution of the σ field along a trajectory across the first-order phase transition boundary (as shown in Fig. 1 traj-A). Fig. 2 shows the time evolution of dynamical cumulants of the σ field along traj-A, with different damping coefficients (solid lines with different colors). The black dotted lines represent the equilibrium cumulants of the σ field at fixed temperature and chemical potential along the trajectory of the evolving heat bath, obtained from the perturbative calculations around the global minimum [1].

As shown in Fig. 2, the magnitude of damping coefficients directly influence the dynamical evolution of σ field. A larger damping coefficient η leads to a slower evolution of the σ field. Besides, the

evolution with a larger damping coefficient tends to develop larger fluctuations at late evolution time. For all three chosen damping coefficients, we find clear memory effects. When the system goes below T_c , the high order cumulants, especially C_3 and C_4 can memorize the sign of the earlier stage during the evolution.

On the other hand, the dynamical cumulants of the σ field are largely enhanced compared with the equilibrium values, after the system evolves across the first order phase transition boundary. As shown in Fig. 1 (middle and right), the effective potential and corresponding probability distribution function shows double well (peak) structures near the first order phase transition above and below T_c . Due to the supercooling effect, the σ field tends to distribute in both minima of the potential, which largely increases the magnitude of high order cumulants.

4 Summary

In this proceeding, we numerically simulate the dynamical evolution of the σ field with a trajectory across the 1st order phase transition boundary, using the event-by-event simulations of the Langevin dynamics. The dynamical cumulants present clear memory effects compared with the equilibrium ones, where C_3 and C_4 can memorize the signs of the early stage during the dynamical evolution. Besides, the dynamical critical fluctuations are largely enhanced after the system across the 1st order phase transition boundary, due to the associated supper cooling effect.

References

- [1] M. A. Stephanov, Phys. Rev. Lett. **102**, 032301 (2009).
- [2] M. A. Stephanov, Phys. Rev. Lett. 107, 052301 (2011).
- [3] M. M. Aggarwal et al. [STAR Collaboration], Phys. Rev. Lett. 105, 022302 (2010).
- [4] L. Adamczyk et al. [STAR Collaboration], Phys. Rev. Lett. 112, 032302 (2014).
- [5] X. Luo [STAR Collaboration], PoS CPOD **2014**, 019 (2014).
- [6] B. Ling and M. A. Stephanov, Phys. Rev. C 93, no. 3, 034915 (2016)
- [7] L. Jiang, P. Li and H. Song, Phys. Rev. C 94, no. 2, 024918 (2016)
- [8] L. Jiang, P. Li and H. Song, Nucl. Phys. A 956, 360 (2016)
- [9] S. Mukherjee, R. Venugopalan and Y. Yin, Phys. Rev. C 92, no. 3, 034912 (2015).
- [10] L. Jiang, S. Wu and H. Song, Nucl. Phys. A **967**, 441 (2017), and paper in preparation.
- [11] D. U. Jungnickel and C. Wetterich, Phys. Rev. D 53, 5142 (1996).
- [12] V. Skokov, B. Friman, E. Nakano, K. Redlich and B.-J. Schaefer, Phys. Rev. D 82, 034029 (2010).
- [13] M. Nahrgang, S. Leupold, C. Herold and M. Bleicher, Phys. Rev. C 84, 024912 (2011).
- [14] P. C. Hohenberg, B. I. Halperin, Rev. Mod. Phys. 49, 435-479 (1977).